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FOR THE NASA-AMES COOLED GRATING
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CORNELL UNIVERSITY

Center for Radiophysics and Space Research

ITHACA, N. Y.

FINAL TECHNICAL REPORT

for

NASA-Ames Research Center

on

Cooperative Agreement NCC 2-79

PROTOTYPE Ge:Ga DETECTORS FOR THE
NASA-AMES COOLED GRATING SPECTROMETER

May 1, 1980 - January 31, 1981

Principal Investigator: J. R. Houck

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FINAL TECHNICAL REPORT

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"PROTOTYPE Ge:Ga DETECTORS FOR THE
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I. INTRODUCTION:

Techniques for the routine fabrication of Ge:Ga detectors were developed under this grant. Numerous detectors were fabricated and tested including seven elements mounted in cavities supplied by E. Erickson of NASA-Ames.

In addition to the usual infrared measurements of responsivity and noise, measurements were made of the detectors response to ionizing radiation. These results are attached as an appendix.

II. DETECTOR PREPARATION:

A. Material:

The detectors were fabricated from a Ge:Ga wafer from Eagle-Pitcher with a room temperature resistivity of $\sim 12\Omega$ cm. The wafer is approximately 2" in diameter and 0.061" thick.

B. Contacts:

The material was ion-implanted with Boron using 10^{14} ions/cm² at 25 Kev and 2×10^{14} ions/cm² at 50 Kev. The crystal was then sputter-cleaned and metalized first with sputtered Ti and then sputter Au.

C. Detector Chips:

The 2x2mm detector chips were cut using an abrasive saw. To remove saw damage the contacts were temporarily protected by a wax film and the sawn surfaces etched with CP-4. The chips were then indium-soldered into the detector cavities.

III. Detector performance:

Once the fabrication techniques were standardized, highly uniform detectors could be quickly produced. Typically, parameters from one detector to the next varied by $\pm 10\%$. In general the characteristics are as follows:

$$R_I \text{ (amp/watt)} \sim 4.0^{\dagger} \text{ (DC)}^x$$

$$V_N (R_L = 5E9) \sim 7.5 \mu v^*$$

$$V_{BIAS} \sim 0.2 \text{ v}$$

[†]Averaged from 40 to 100 μm .

^xOne can expect to achieve about 75% of the above responsivity at 20Hz.

^{*}Background power of 5E-11 watts.

APPENDIX

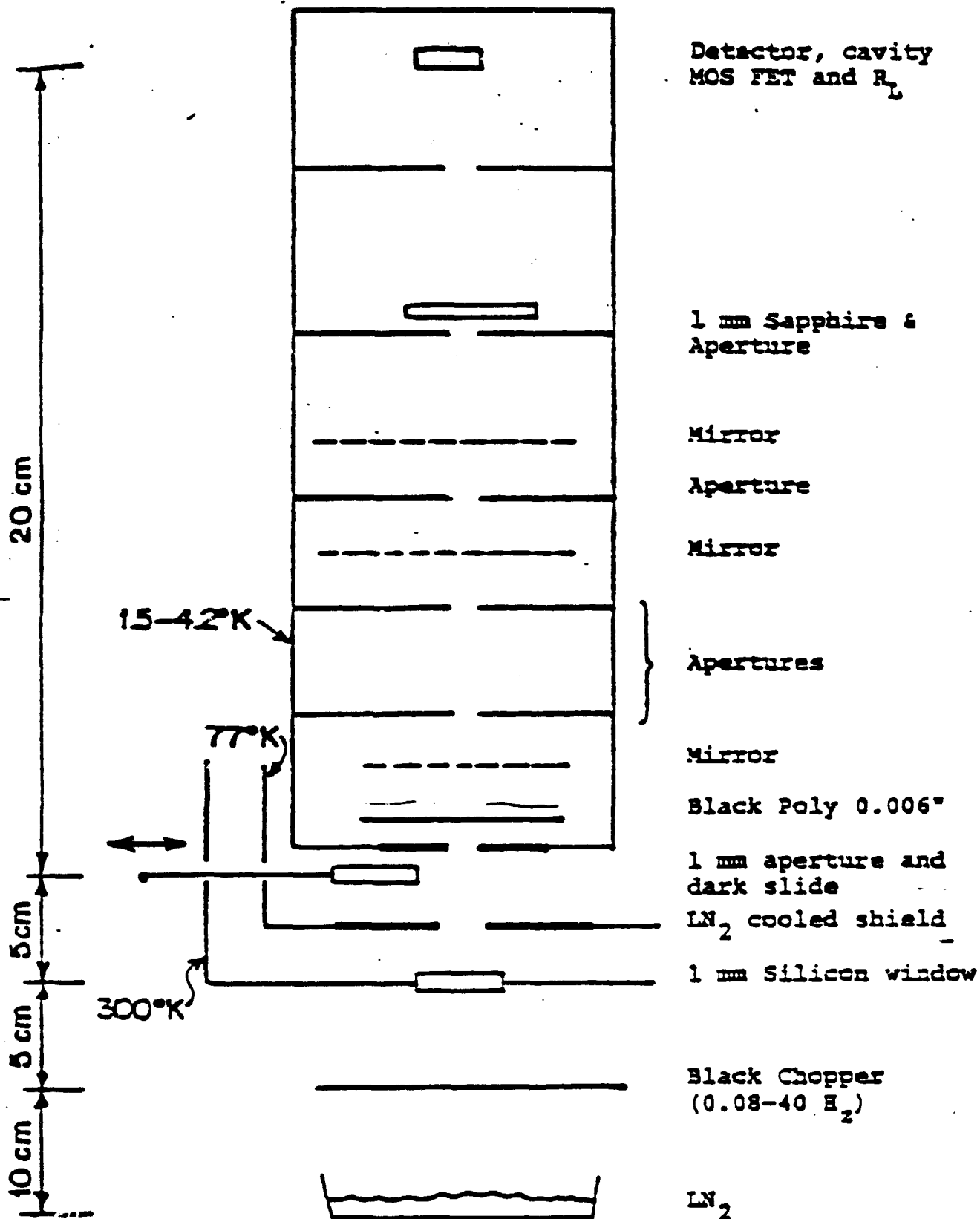
TEST CONDITIONS

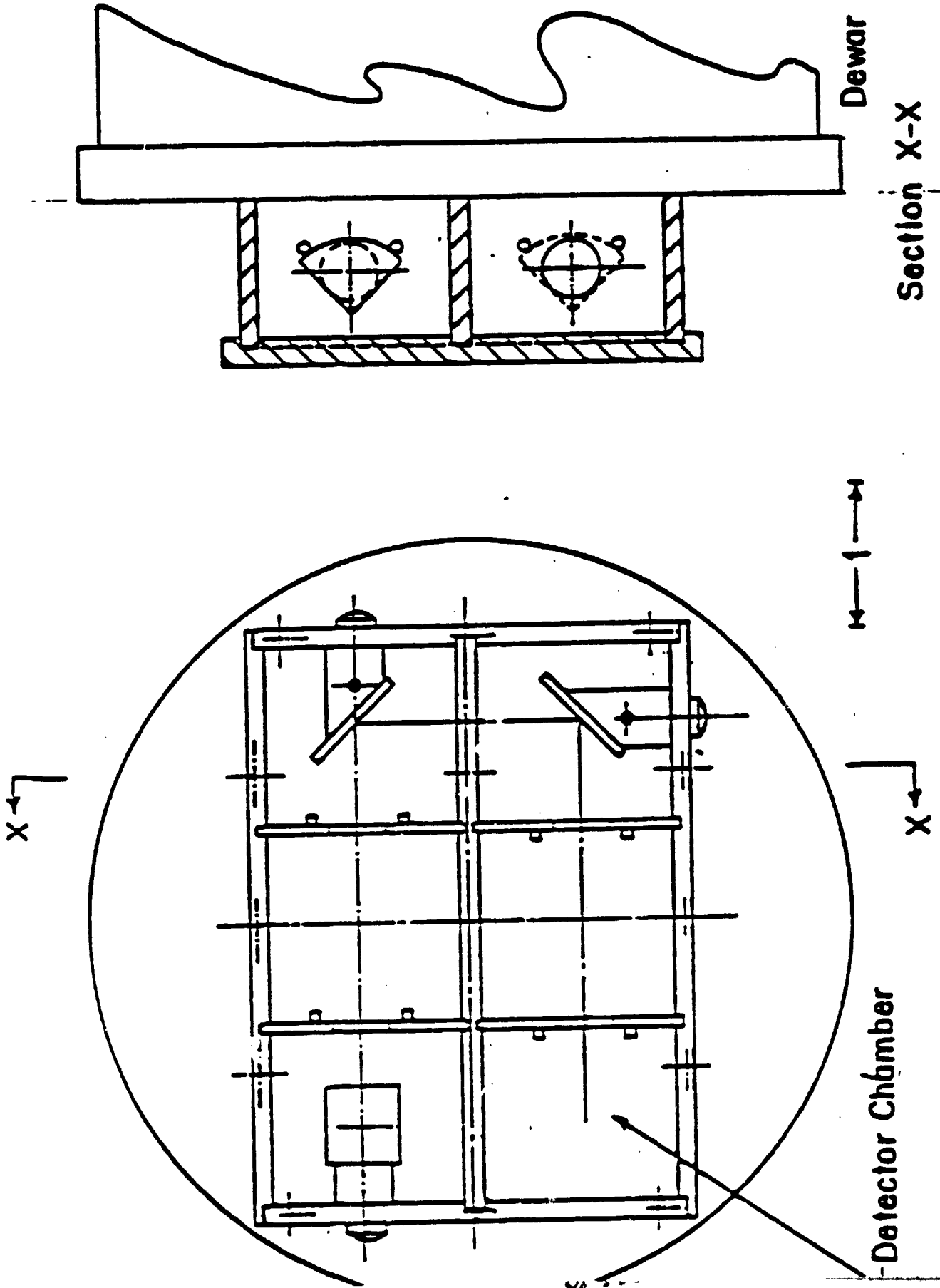
- (1) Detector: Ge:Ga $2 \times 2 \times 1\frac{1}{2}$ mm
Eagle-Picher; LC 1606 with ion implanted contacts.
- (2) Cavity: Gold-plated copper $6 \text{ mm } \phi \times 1\frac{3}{4} \text{ mm}$ deep entrance hole $1.34 \text{ (mm}^2\text{)}$



- (3) Led: Standard red led with heat sink
Typical operation 0.5 ma
No detectable heat up at this current
- (4) Optical System: OPI (see next two pages)
Signal power levels:
Low level: $1.5 \times 10^{-13} \text{ w}$ in band (with 2% filter)
High level: $5 \times 10^{-11} \text{ w}$ in band
Dark (slide closed $P_B \ll 10^{-14} \text{ w}$)
- (5) Signal Source: Chopped LN_2 "Black Body"
- (6) Electronics:
Eltec load resistor
 $R(2^\circ\text{K}) = 5 \times 10^3 \Omega$
Standard TIA with
Balanced JFET ($\sim 100^\circ\text{K}$) 2N6484

OPTICAL PATH (unfolded) OPI CORNELL FILTERS





TEST RESULTS

(1) Responsivity vs. LED-Induced Background:

The AC (20 Hz) and DC responsivities increased by 20x for a LED-induced photo current of 1.4×10^{-9} Amps ($V_B = 250\text{ mv}$) (The DC responsivity changed by less than 30% for an IR-induced photo current of the same magnitude.)

(2) Recovery Time:

The recovery time (the time required for the excess (LED induced) responsivity to decay to $1/e$) depends on many factors. These include temperature, IR background and bias voltage.

<u>Test No.</u>	<u>T (Recovery)</u>	<u>Condition</u>
1	~36 minutes	Dark $P_B \ll 10^{-14}\text{w}$
2	"	Very low bkg; $P_B \sim 1.5 \times 10^{-14}\text{w}$
3	"	Low bkg; $P_B \sim 1.5 \times 10^{-13}\text{w}$
4	6.1 minutes	High bkg; $P_B \sim 5 \times 10^{-11}\text{w}$ (Reduced bias 105 mv)
5	4 minutes	High bkg; $P_B \sim 5 \times 10^{-11}\text{w}$ (Normal bias 200 mv)
6	6.3 minutes	Low bkg; $1.5 \times 10^{-13}\text{w}$ 3.1°K
7	~20 sec	$V_B = 1.0\text{ v}$ - full breakdown

Tests 1,2,3,4,5 and 7: $T_{\text{DET.}} \approx 1.7^\circ\text{K}$ ($T_{\text{BATH}} = 1.6^\circ\text{K}$)

Conclusions:

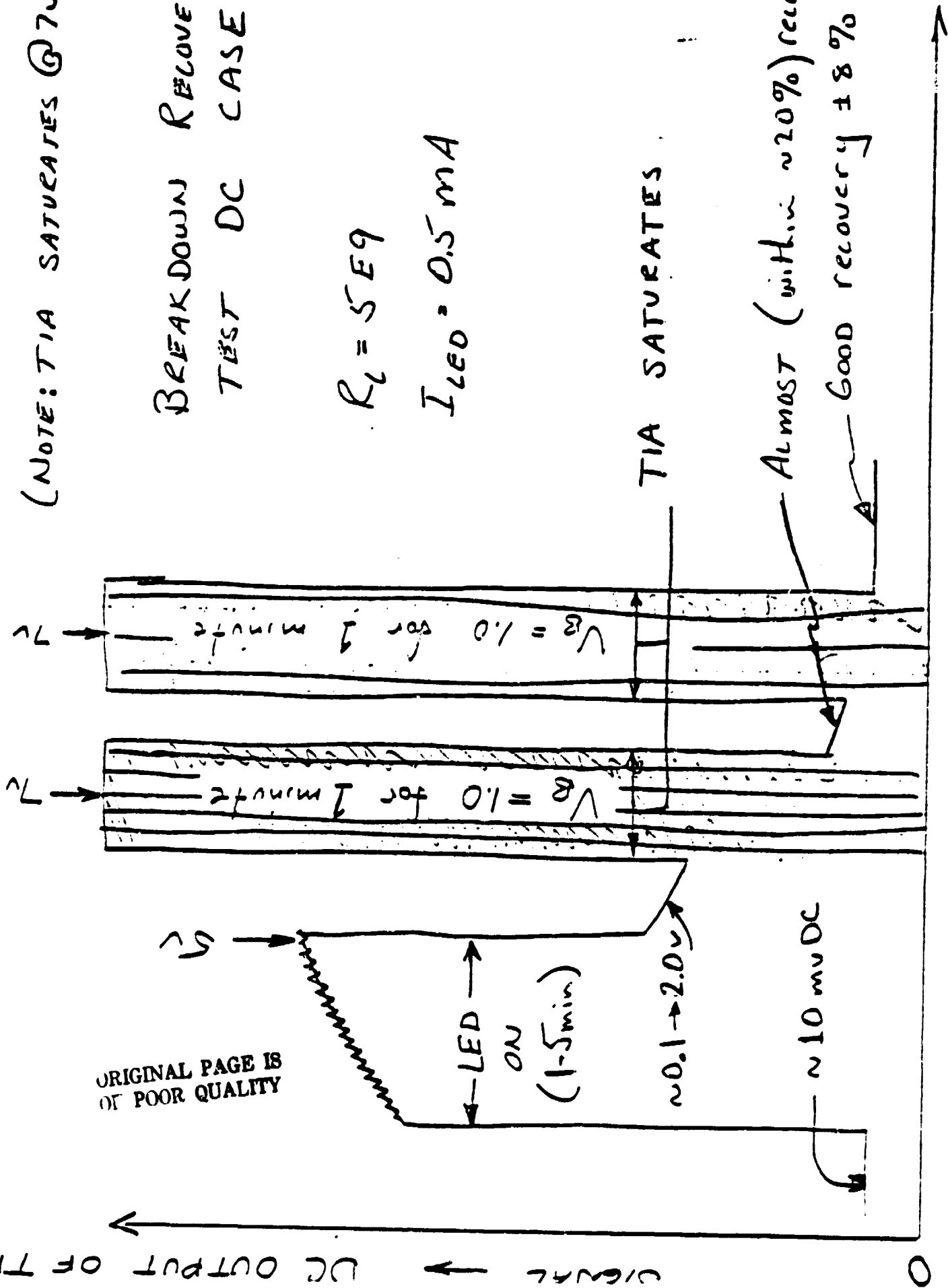
- (1) LED-induced-enhanced responsivity shows the same characteristics as γ -induced responsivity.
- (2) Under low-background, low-temperature conditions the recovery time is very long, ~ hours.
- (3) Speedy recovery can be achieved by driving the detector into avalanche breakdown by increasing the bias.

(NOTE: TIA SATURATES @ 7V)

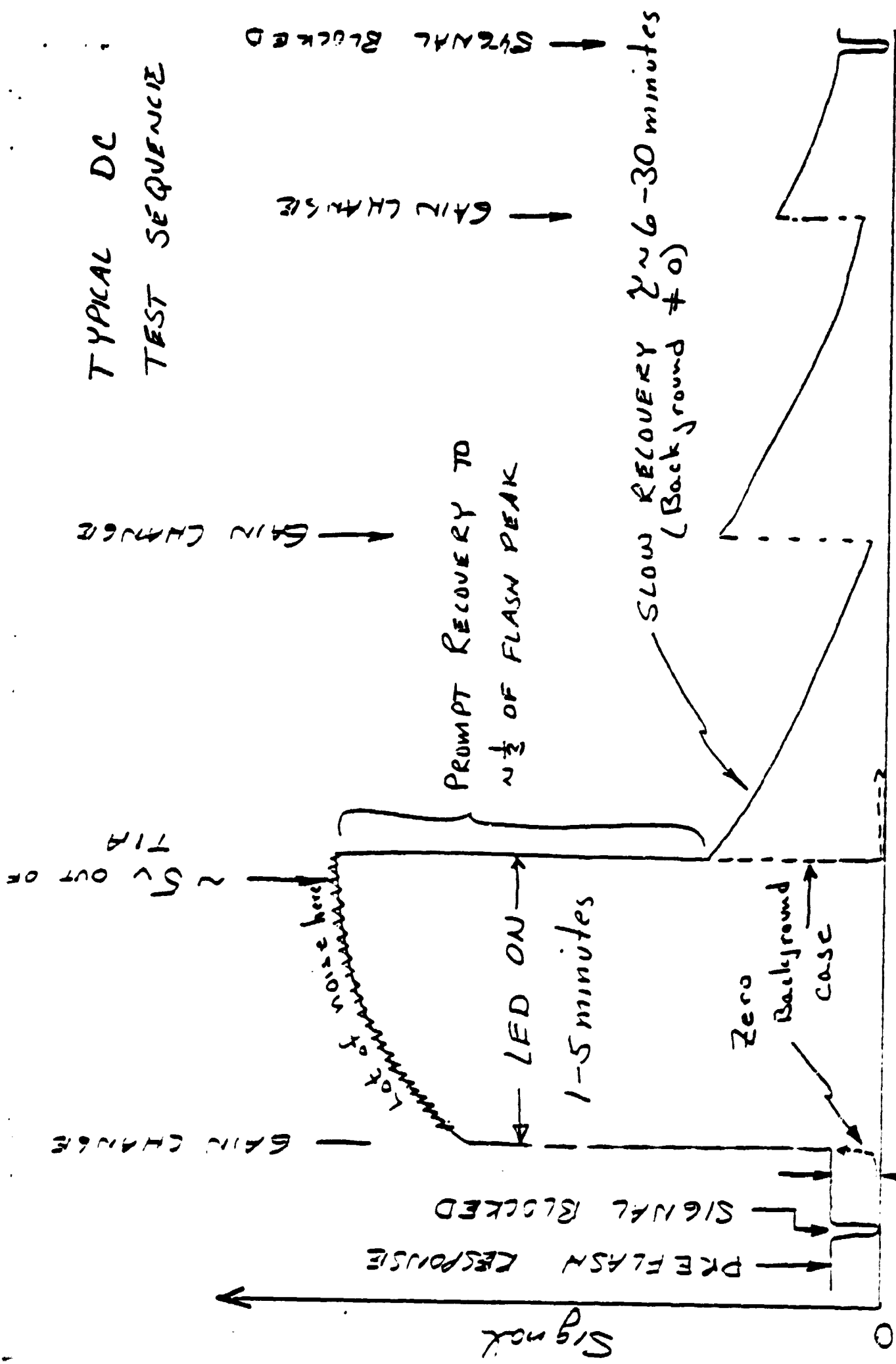
BREAKDOWN RECOVERY
TEST DC CASE

$$R_L = 5 \text{ E9}$$

$$I_{LED} = 0.5 \text{ mA}$$



time →



TYPICAL DC
TEST SEQUENCE